

Simplified P_N Equations for Nonclassical Transport with Isotropic Scattering

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I Introduction

The nonclassical transport equation [1] was developed to address transport problems in which the particle flux is not attenuated exponentially. It consists of a linear Boltzmann equation on an extended phase space, able to model particle transport for any given free-path distribution.

In this paper we consider the one-speed nonclassical transport equation with isotropic scattering. This equation is written as

$$\frac{\partial \Psi}{\partial s}(s) + \boldsymbol{\Omega} \cdot \nabla \Psi(s) + \Sigma_t(s) \Psi(s) = \frac{\delta(s)}{4\pi} \left[\int_{4\pi} \int_0^\infty c \Sigma_t(s') \Psi(\mathbf{x}, \boldsymbol{\Omega}', s') ds' d\Omega' + Q(\mathbf{x}) \right], \quad (1)$$

where s describes the free-path of a particle (distance traveled since the particle's previous interaction), $\Psi(s) = \Psi(\mathbf{x}, \boldsymbol{\Omega}, s)$ is the nonclassical angular flux, c is the scattering ratio (probability of scattering), and Q is an isotropic source. The classical scalar particle flux $\Phi(\mathbf{x})$ can be recovered from the solution of Eq. (1) by integrating over the free-path s and over angle: $\Phi(\mathbf{x}) = \int_{4\pi} \int_0^\infty \Psi(\mathbf{x}, \boldsymbol{\Omega}, s) ds d\Omega$.

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The total cross section Σ_t is a function of the free-path s and satisfies $p(s) = \Sigma_t(s)e^{-\int_0^s \Sigma_t(s')ds'}$, where $p(s)$ is the free-path distribution function. For $m = 0, 1, 2, \dots$, we define the m -th raw moment of $p(s)$ as $\langle s^m \rangle = \int_0^\infty s^m p(s) ds$. The following identity holds for $m = 1, 2, \dots$:

$$\langle s^m \rangle = m \int_0^\infty s^{m-1} e^{-\int_0^s \Sigma_t(s')ds'} ds. \quad (2)$$

A nonclassical diffusion approximation of Eq. (1) has been previously derived [1]. In this paper we use an asymptotic analysis to derive more accurate diffusion approximations to Eq. (1). If $p(s)$ is given by an exponential (classical transport), these approximations reduce to the simplified P_N (SP_N) equations [2]; therefore, they are labeled *nonclassical* SP_N equations.

II Asymptotic Analysis

Let us write Eq. (1) in the mathematically equivalent form

$$\begin{aligned} \frac{\partial \Psi}{\partial s}(s) + \boldsymbol{\Omega} \cdot \nabla \Psi(s) + \Sigma_t(s)\Psi(s) &= 0, \quad s > 0, \\ \Psi(0) &= \frac{1}{4\pi} \left[\int_{4\pi} \int_0^\infty c \Sigma_t(s') \Psi(\mathbf{x}, \boldsymbol{\Omega}', s') ds' d\Omega' + Q(\mathbf{x}) \right], \end{aligned}$$

where $\Psi(0) = \lim_{s \rightarrow 0^+} \Psi(s) = \Psi(0^+)$. Defining $0 < \varepsilon \ll 1$, we perform the following scaling: $\Sigma_t(s) = \varepsilon^{-1} \Sigma_t(s/\varepsilon)$; $c = 1 - \varepsilon^2 \kappa$; and $Q(\mathbf{x}) = \varepsilon q(\mathbf{x})$; where κ and q are $O(1)$.

Under this scaling,

$$\langle s^m \rangle = \varepsilon^m \int_0^\infty s^m \Sigma_t(s) e^{-\int_0^s \Sigma_t(s')ds'} ds = \varepsilon^m \langle s^m \rangle_\varepsilon, \quad (3)$$

where $\langle s^m \rangle_\varepsilon$ is $O(1)$.

Next, we define $\psi(\mathbf{x}, \mathbf{\Omega}, s) \equiv \frac{\varepsilon \langle s \rangle_\epsilon}{e^{-\int_0^s \Sigma_t(s') ds'}} \Psi(\mathbf{x}, \mathbf{\Omega}, \varepsilon s)$. This satisfies

$$\frac{\partial \psi}{\partial s}(s) + \varepsilon \mathbf{\Omega} \cdot \nabla \psi(s) = 0, \quad s > 0, \quad (4a)$$

$$\psi(0) = \frac{1}{4\pi} \left[\int_{4\pi} \int_0^\infty (1 - \varepsilon^2 \kappa) p(s') \psi(\mathbf{x}, \mathbf{\Omega}', s') ds' d\Omega' + \varepsilon^2 \langle s \rangle_\epsilon q(\mathbf{x}) \right], \quad (4b)$$

and the classical scalar flux can be written as $\Phi(\mathbf{x}) = \int_{4\pi} \int_0^\infty \psi(\mathbf{x}, \mathbf{\Omega}, s) \frac{e^{-\int_0^s \Sigma_t(s') ds'}}{\langle s \rangle_\epsilon} ds d\Omega$.

Integrating Eq. (4a) over $0 < s' < s$ and using Eq. (4b), we obtain

$$\left(I + \varepsilon \mathbf{\Omega} \cdot \nabla \int_0^s (\cdot) ds \right) \psi = \frac{1}{4\pi} \left[\int_0^\infty (1 - \varepsilon^2 \kappa) p(s') \varphi(\mathbf{x}, s') ds' + \varepsilon^2 \langle s \rangle_\epsilon q \right],$$

where $\varphi(\mathbf{x}, s) = \int_{4\pi} \psi(\mathbf{x}, \mathbf{\Omega}, s) d\Omega$. Inverting the operator on the left-hand side of the above equation and expanding it in a power series, we obtain

$$\psi = \left(\sum_{n=0}^\infty (-\varepsilon)^n \left(\mathbf{\Omega} \cdot \nabla \int_0^s (\cdot) ds \right)^n \right) \left[\int_0^\infty \frac{1 - \varepsilon^2 \kappa}{4\pi} p(s') \varphi(\mathbf{x}, s') ds' + \varepsilon^2 \langle s \rangle_\epsilon \frac{q}{4\pi} \right]. \quad (5)$$

Let us define $\nabla_0 = \frac{1}{3} \nabla^2$ and $\mathcal{B} = \nabla_0 (\int_0^s (\cdot) ds)^2$. Then, using the identity [3]

$$\frac{1}{4\pi} \int_{4\pi} \left(\mathbf{\Omega} \cdot \nabla \int_0^s (\cdot) ds \right)^n d\Omega = \frac{1 + (-1)^n}{2} \frac{3^{n/2}}{n+1} \mathcal{B}^{n/2},$$

for $n = 0, 1, 2, \dots$, we integrate Eq. (5) over the unit sphere and obtain

$$\varphi = \left(\sum_{n=0}^\infty \frac{\varepsilon^{2n}}{2n+1} (3\mathcal{B})^n \right) \left[\int_0^\infty (1 - \varepsilon^2 \kappa) p(s') \varphi(\mathbf{x}, s') ds' + \varepsilon^2 \langle s \rangle_\epsilon q \right].$$

Inverting the operator on the right-hand side of this equation and once again expand-

ing it in a power series, we get

$$\left(I - \varepsilon^2 \mathcal{B} - \frac{4\varepsilon^4}{5} \mathcal{B}^2 - \frac{44\varepsilon^6}{35} \mathcal{B}^3 + O(\varepsilon^8)\right) \varphi = \int_0^\infty (1 - \varepsilon^2 \kappa) p(s') \varphi(\mathbf{x}, s') ds' + \varepsilon^2 \langle s \rangle_\epsilon q. \quad (6a)$$

The solution of this equation is

$$\varphi(\mathbf{x}, s) = \left(I + \varepsilon^2 \frac{s^2}{2!} \nabla_0 + \frac{9\varepsilon^4}{5} \frac{s^4}{4!} \nabla_0^2 + \frac{27\varepsilon^6}{7} \frac{s^6}{6!} \nabla_0^3 + O(\varepsilon^8)\right) \phi(\mathbf{x}), \quad (6b)$$

where $\phi(\mathbf{x}) = \sum_{n=0}^\infty \varepsilon^2 \phi_{2n}(\mathbf{x})$, with $\phi_{2n}(\mathbf{x})$ undetermined at this point.

We now multiply Eqs. (6) by $e^{-\int_0^s \Sigma_t(s') ds'} / \langle s \rangle_\epsilon$ and operate on them by $\int_0^\infty (\cdot) ds$.

Using Eqs. (2), (3) and (5), Eq. (6b) yields an expression for the scalar flux:

$$\Phi(\mathbf{x}) = \left(I + \varepsilon^2 \frac{\langle s^3 \rangle_\epsilon}{3! \langle s \rangle_\epsilon} \nabla_0 + \frac{9\varepsilon^4}{5} \frac{\langle s^5 \rangle_\epsilon}{5! \langle s \rangle_\epsilon} \nabla_0^2 + \frac{27\varepsilon^6}{7} \frac{\langle s^7 \rangle_\epsilon}{7! \langle s \rangle_\epsilon} \nabla_0^3 + O(\varepsilon^8)\right) \phi(\mathbf{x}).$$

We can write

$$\int_0^\infty p(s) \varphi(\mathbf{x}, s) ds = \left(\sum_{n=0}^\infty \varepsilon^{2n} U_n \nabla_0^n\right) \Phi(\mathbf{x}),$$

with

$$\begin{aligned}
U_0 &= 1; \\
U_1 &= \frac{\langle s^2 \rangle_\epsilon}{2!} - \frac{\langle s^3 \rangle_\epsilon}{3! \langle s \rangle_\epsilon}; \\
U_2 &= \frac{9}{5} \left[\frac{\langle s^4 \rangle_\epsilon}{4!} - \frac{\langle s^5 \rangle_\epsilon}{5! \langle s \rangle_\epsilon} \right] - \frac{\langle s^3 \rangle_\epsilon}{3! \langle s \rangle_\epsilon} U_1; \\
U_3 &= \frac{27}{7} \left[\frac{\langle s^6 \rangle_\epsilon}{6!} - \frac{\langle s^7 \rangle_\epsilon}{7! \langle s \rangle_\epsilon} \right] - \frac{9}{5} \frac{\langle s^5 \rangle_\epsilon}{5! \langle s \rangle_\epsilon} U_1 - \frac{\langle s^3 \rangle_\epsilon}{3! \langle s \rangle_\epsilon} U_2; \\
&\dots
\end{aligned}$$

Equation (6a) can be rewritten as

$$\left(\sum_{n=0}^{\infty} \varepsilon^{2n} V_n \nabla_0^n \right) \Phi(\mathbf{x}) = (1 - \varepsilon^2 \kappa) \left(\sum_{n=0}^{\infty} \varepsilon^{2n} U_n \nabla_0^n \right) \Phi(\mathbf{x}) + \varepsilon^2 \langle s \rangle_\epsilon q(\mathbf{x}), \quad (7)$$

where

$$\begin{aligned}
V_0 &= 1; \\
V_1 &= -\frac{\langle s^3 \rangle_\epsilon}{3! \langle s \rangle_\epsilon} V_0; \\
V_2 &= -\frac{9}{5} \frac{\langle s^5 \rangle_\epsilon}{5! \langle s \rangle_\epsilon} V_0 - \frac{\langle s^3 \rangle_\epsilon}{3! \langle s \rangle_\epsilon} V_1; \\
V_3 &= -\frac{27}{7} \frac{\langle s^7 \rangle_\epsilon}{7! \langle s \rangle_\epsilon} V_0 - \frac{9}{5} \frac{\langle s^5 \rangle_\epsilon}{5! \langle s \rangle_\epsilon} V_1 - \frac{\langle s^3 \rangle_\epsilon}{3! \langle s \rangle_\epsilon} V_2; \\
&\dots
\end{aligned}$$

Finally, rearranging the terms in Eq. (7) we get

$$\left(\sum_{n=0}^{\infty} \varepsilon^{2n} \left[W_{n+1} \nabla_0^{n+1} + \kappa U_n \nabla_0^n \right] \right) \Phi(\mathbf{x}) = \langle s \rangle_{\epsilon} q(\mathbf{x}), \quad (8)$$

where $W_n = V_n - U_n$. If we discard the terms of $O(\varepsilon^{2n})$ in this equation, we obtain a partial differential equation for $\Phi(\mathbf{x})$ of order $2n$. We will use this approach to explicitly derive the nonclassical SP₁, SP₂, and SP₃ equations. Higher-order equations can be derived from Eq. (8) by continuing to follow the same procedure.

II.A Nonclassical diffusion equation (nonclassical SP₁)

We discard the terms of $O(\varepsilon^2)$ in Eq. (8) and rewrite the equation as

$$W_1 \nabla_0 \Phi(\mathbf{x}) + \kappa \Phi(\mathbf{x}) = \langle s \rangle_{\epsilon} q(\mathbf{x}).$$

Multiplying this equation by ε and reverting to the original unscaled parameters, we obtain

$$-\frac{1}{6} \frac{\langle s^2 \rangle}{\langle s \rangle} \nabla^2 \Phi(\mathbf{x}) + \frac{1-c}{\langle s \rangle} \Phi(\mathbf{x}) = Q(\mathbf{x}), \quad (9)$$

which is the nonclassical diffusion equation as derived in [1].

II.B Nonclassical simplified P₂ equation

Discarding the terms of $O(\varepsilon^4)$ in Eq. (8) and rearranging the terms, we get

$$-\left(I + \varepsilon^2 \frac{W_2 \nabla_0 + \kappa U_1}{W_1} \right) W_1 \nabla_0 \Phi(\mathbf{x}) = \kappa \Phi(\mathbf{x}) - \langle s \rangle_{\epsilon} q(\mathbf{x}).$$

Operating on this equation by $(I - \varepsilon^2 [W_2 \nabla_0 + \kappa U_1] / W_1)$ and discarding terms of $O(\varepsilon^4)$, it becomes

$$W_1 \nabla_0 \left[\Phi(\mathbf{x}) - \varepsilon^2 \frac{W_2}{W_1^2} \left[\kappa \Phi(\mathbf{x}) - \langle s \rangle_\epsilon q(\mathbf{x}) \right] \right] + \kappa \left[1 - \varepsilon^2 \kappa \frac{U_1}{W_1} \right] \Phi(\mathbf{x}) = \left[1 - \varepsilon^2 \kappa \frac{U_1}{W_1} \right] \langle s \rangle_\epsilon q(\mathbf{x}).$$

Finally, we multiply this equation by ε and revert to the original unscaled parameters to obtain the nonclassical SP₂ equation

$$-\frac{1}{6} \frac{\langle s^2 \rangle}{\langle s \rangle} \nabla^2 \left[\Phi(\mathbf{x}) + \lambda_1 \left[(1 - c) \Phi(\mathbf{x}) - \langle s \rangle Q(\mathbf{x}) \right] \right] + \frac{1 - c}{\langle s \rangle} \left[1 - \beta_1 (1 - c) \right] \Phi(\mathbf{x}) = \left[1 - \beta_1 (1 - c) \right] Q(\mathbf{x}), \quad (10)$$

$$\text{with } \lambda_1 = \frac{3}{10} \frac{\langle s^4 \rangle}{\langle s^2 \rangle^2} - \frac{1}{3} \frac{\langle s^3 \rangle}{\langle s \rangle \langle s^2 \rangle} \text{ and } \beta_1 = \frac{1}{3} \frac{\langle s^3 \rangle}{\langle s \rangle \langle s^2 \rangle} - 1.$$

II.C Nonclassical simplified P₃ equations

Discarding the terms of $O(\varepsilon^6)$ in Eq. (8), we have

$$\left(W_1 \nabla_0 + \varepsilon^2 [W_2 \nabla_0^2 + \kappa U_1 \nabla_0] + \varepsilon^4 [W_3 \nabla_0^3 + \kappa U_2 \nabla_0^2] \right) \Phi(\mathbf{x}) + \kappa \Phi(\mathbf{x}) = \langle s \rangle_\epsilon q(\mathbf{x}).$$

We define $\nu(\mathbf{x}) = \left(I + \varepsilon^2 \frac{W_3 \nabla_0 + \kappa U_2}{W_2} \right) \frac{\varepsilon^2}{2} \frac{W_2}{W_1} \nabla_0 \Phi(\mathbf{x})$, and rewrite this equation as

$$W_1 \nabla_0 \left[\Phi(\mathbf{x}) + 2\nu(\mathbf{x}) + \varepsilon^2 \kappa \frac{U_1}{W_1} \Phi(\mathbf{x}) \right] + \kappa \Phi(\mathbf{x}) = \langle s \rangle_\epsilon q(\mathbf{x}).$$

Operating on $\nu(\mathbf{x})$ by $(I - \varepsilon^2[W_3\nabla_0 + \kappa U_2]/W_2)$ and discarding terms of $O(\varepsilon^6)$, we get

$$\varepsilon^2 \frac{W_1}{\langle s \rangle_\epsilon} \nabla_0 \left[\frac{W_3}{W_1 W_2} \nu(\mathbf{x}) + \frac{1}{2} \frac{W_2}{W_1^2} \Phi(\mathbf{x}) \right] + \frac{1}{\langle s \rangle_\epsilon} \left[1 - \varepsilon^2 \kappa \frac{U_2}{W_2} \right] \nu(\mathbf{x}) = 0.$$

Finally, we multiply the two equations above by ε and $1/\varepsilon$, respectively, and revert to the original unscaled parameters to obtain the nonclassical SP₃ equations

$$-\frac{1}{6} \frac{\langle s^2 \rangle}{\langle s \rangle} \nabla^2 \left[[1 + \beta_1(1 - c)] \Phi(\mathbf{x}) + 2\nu(\mathbf{x}) \right] + \frac{1 - c}{\langle s \rangle} \Phi(\mathbf{x}) = Q(\mathbf{x}), \quad (11a)$$

$$-\frac{1}{6} \frac{\langle s^2 \rangle}{\langle s \rangle} \nabla^2 \left[\frac{\lambda_1}{2} \Phi(\mathbf{x}) + \lambda_2 \nu(\mathbf{x}) \right] + \frac{1 - \beta_2(1 - c)}{\langle s \rangle} \nu(\mathbf{x}) = 0, \quad (11b)$$

where λ_1 and β_1 are the same as in the SP₂ equations, and

$$\lambda_2 = \frac{1}{10\langle s^2 \rangle \langle s^3 \rangle - 9\langle s \rangle \langle s^4 \rangle} \left[\frac{9}{5} \langle s^5 \rangle - \frac{27}{21} \frac{\langle s \rangle \langle s^6 \rangle}{\langle s^2 \rangle} + 3 \frac{\langle s^3 \rangle \langle s^4 \rangle}{\langle s^2 \rangle} - \frac{10}{3} \frac{\langle s^3 \rangle^2}{\langle s \rangle} \right],$$

$$\beta_2 = \frac{1}{10\langle s^2 \rangle_\epsilon \langle s^3 \rangle_\epsilon - 9\langle s \rangle_\epsilon \langle s^4 \rangle_\epsilon} \left[\frac{10}{3} \frac{\langle s^3 \rangle_\epsilon^2}{\langle s \rangle_\epsilon} - \frac{9}{5} \langle s^5 \rangle_\epsilon \right] - 1.$$

III Classical SP_N Equations

If $\Sigma_t(s) = \Sigma_t$ is independent of s , the nonclassical SP_N equations derived in the previous section reduce to the SP_N approximations to the classical transport equation [2]. Under this assumption, the free-path distribution $p(s)$ is an exponential and $\langle s^m \rangle = m! \Sigma_t^{-m}$.

Introducing this result into the nonclassical diffusion approximation given by Eq. (9), one can easily see that it reduces to the classical diffusion equation

$$-\frac{1}{3\Sigma_t} \nabla^2 \Phi(\mathbf{x}) + \Sigma_a \Phi(\mathbf{x}) = Q(\mathbf{x}).$$

We also obtain $\lambda_1 = \frac{4}{5}$; $\lambda_2 = \frac{11}{7}$; and $\beta_1 = \beta_2 = 0$. In this case, the nonclassical SP₂ equation (10) reduces to

$$-\frac{1}{3\Sigma_t}\nabla^2\left[\Phi(\mathbf{x}) + \frac{4}{5}\frac{\Sigma_a\Phi(\mathbf{x}) - Q(\mathbf{x})}{\Sigma_t}\right] + \Sigma_a\Phi(\mathbf{x}) = Q(\mathbf{x}),$$

which is the classical SP₂ equation. The nonclassical SP₃ equations (11) reduce to

$$\begin{aligned} -\frac{1}{3\Sigma_t}\nabla^2\left[\Phi(\mathbf{x}) + 2\nu(\mathbf{x})\right] + \Sigma_a\Phi(\mathbf{x}) &= Q(\mathbf{x}), \\ -\frac{1}{3\Sigma_t}\nabla^2\left[\frac{2}{5}\Phi(\mathbf{x}) + \frac{11}{7}\nu(\mathbf{x})\right] + \Sigma_t\nu(\mathbf{x}) &= 0, \end{aligned}$$

which are the classical SP₃ equations.

Moreover, $U_1 = U_2 = U_3 \dots = 0$ and $V_0 = 1$; $V_1 = -\frac{1}{\sigma_t^2}$; $V_2 = -\frac{4}{5\sigma_t^4}$; $V_3 = -\frac{44}{35\sigma_t^6}$; with $\sigma_t = \Sigma_t/\varepsilon$. Thus, Eq. (8) becomes

$$-\left(\frac{1}{\sigma_t}\nabla_0 + \frac{4\varepsilon^2}{5\sigma_t^3}\nabla_0^2 + \frac{44\varepsilon^4}{35\sigma_t^5}\nabla_0^3 + O(\varepsilon^6)\right)\Phi(\mathbf{x}) + \kappa\Phi(\mathbf{x}) = q(\mathbf{x}).$$

This is the asymptotic approximation to the classical transport equation with isotropic scattering that can be used to obtain the classical SP_N equations [2].

III.A A comment on boundary conditions

Our asymptotic analysis does not yield boundary conditions. Therefore, in order to generate the numerical results in this paper, we manipulate the nonclassical SP_N equations to a classical form with modified parameters, and then use classical (Marshak) vacuum boundary conditions.

Specifically, the nonclassical SP₁ equation with vacuum boundary conditions is

written as

$$\begin{aligned} -\frac{1}{3\widehat{\Sigma}_t}\nabla^2\Phi(\mathbf{x}) + \widehat{\Sigma}_a\Phi(\mathbf{x}) &= Q(\mathbf{x}), \\ \frac{1}{2}\Phi(\mathbf{x}) - \frac{1}{3\widehat{\Sigma}_t}\mathbf{n} \cdot \nabla\Phi(\mathbf{x}) &= 0, \end{aligned}$$

with

$$\widehat{\Sigma}_t = 2\frac{\langle s \rangle}{\langle s^2 \rangle}, \quad \widehat{\Sigma}_a = \frac{1-c}{\langle s \rangle}.$$

Similarly, the nonclassical SP₂ equation with vacuum boundary conditions becomes

$$\begin{aligned} -\frac{1}{3\widehat{\Sigma}_t}\nabla^2\widehat{\Phi}(\mathbf{x}) + \widehat{\Sigma}_a\widehat{\Phi}(\mathbf{x}) &= \widehat{Q}(\mathbf{x}), \\ \frac{1}{2}\widehat{\Phi}(\mathbf{x}) - \frac{1}{3\widehat{\Sigma}_t}\mathbf{n} \cdot \nabla\widehat{\Phi}(\mathbf{x}) &= 0, \end{aligned}$$

with

$$\begin{aligned} \widehat{\Sigma}_t &= 2\frac{\langle s \rangle}{\langle s^2 \rangle}, & \widehat{\Sigma}_a &= \frac{(1-c)}{\langle s \rangle} \frac{1-\beta_1(1-c)}{1+\lambda_1(1-c)}, \\ \widehat{Q}(\mathbf{x}) &= \frac{1-\beta_1(1-c)}{1+\lambda_1(1-c)}Q(\mathbf{x}), & \Phi(\mathbf{x}) &= \frac{\widehat{\Phi}(\mathbf{x}) + \lambda_1\langle s \rangle Q(\mathbf{x})}{1+\lambda_1(1-c)}. \end{aligned}$$

Finally, the nonclassical SP₃ equations with vacuum boundary conditions become

$$\begin{aligned} -\frac{1}{3\widehat{\Sigma}_t}\nabla^2[\Phi(\mathbf{x}) + 2\widehat{\Phi}_2(\mathbf{x})] + \widehat{\Sigma}_a\Phi(\mathbf{x}) &= \widehat{Q}(\mathbf{x}), \\ -\frac{1}{3\widehat{\Sigma}_t}\nabla^2\left[\frac{2}{5}\Phi(\mathbf{x}) + \left(\frac{4}{5} + \frac{27\widehat{\Sigma}_t}{35\widehat{\Sigma}_3}\right)\widehat{\Phi}_2(\mathbf{x})\right] + \widehat{\Sigma}_2\widehat{\Phi}_2(\mathbf{x}) &= 0, \\ \frac{1}{2}\Phi(\mathbf{x}) - \frac{1}{3\widehat{\Sigma}_t}\mathbf{n} \cdot \nabla\Phi(\mathbf{x}) - \frac{2}{3\widehat{\Sigma}_t}\mathbf{n} \cdot \nabla\widehat{\Phi}_2(\mathbf{x}) + \frac{5}{8}\widehat{\Phi}_2(\mathbf{x}) &= 0, \\ -\frac{1}{8}\Phi(\mathbf{x}) + \frac{5}{8}\widehat{\Phi}_2(\mathbf{x}) - \frac{3}{7\widehat{\Sigma}_3}\mathbf{n} \cdot \nabla\widehat{\Phi}_2(\mathbf{x}) &= 0, \end{aligned}$$

with

$$\begin{aligned}
\hat{\Phi}_2(\mathbf{x}) &= \frac{\nu(\mathbf{x})}{1 + \beta_1(1 - c)}, & \hat{\Sigma}_t &= 2 \frac{\langle s \rangle}{\langle s^2 \rangle}, \\
\hat{\Sigma}_a &= \frac{(1 - c)}{\langle s \rangle} \frac{1}{1 + \beta_1(1 - c)}, & \hat{Q}(\mathbf{x}) &= \frac{Q(\mathbf{x})}{1 + \beta_1(1 - c)}, \\
\hat{\Sigma}_2 &= \frac{4 [1 + \beta_1(1 - c)] [1 - \beta_2(1 - c)]}{5 \lambda_1 \langle s \rangle}, & \hat{\Sigma}_3 &= \frac{27}{28} \frac{\lambda_1 \hat{\Sigma}_t}{\lambda_2 [1 + \beta_1(1 - c)] - \lambda_1}.
\end{aligned}$$

IV Numerical Results

We consider transport taking place in a 1-D random periodic system similar to the one introduced in [4], consisting of a random segment of periodically arranged solid and void layers of equal width $\ell = 0.5$. The total width of the system is given by $2X = 4\ell M$, where the integer M (the total length of each material in the system) satisfies $M = \varepsilon^{-1}$. Vacuum boundary conditions are assigned at $x = \pm X$.

Cross sections and source in the void material are 0; the parameters of the solid material are $\Sigma_{t1} = 1$ and $Q_1 = 0.2/M^2$. The absorption ratio is given by $1 - c = 0.1/M^2$. As M increases, ε decreases, and the 1-D system approaches the diffusive limit described in the asymptotic analysis.

To generate benchmark results for comparison, we used the same procedure presented in [5]. In this procedure, we obtain a physical realization of the system by choosing a continuous segment of two full layers (one of each material) and randomly placing the coordinate $x = 0$ in this segment. Given this fixed realization of the system, the cross sections and source are now deterministic functions of space.

We solve the transport equation numerically for this realization using (i) the standard discrete ordinate method with a 16-point Gauss-Legendre quadrature set (S_{16}); and (ii) diamond differencing for the spatial discretization. This procedure is repeated for different realizations of the random system. Finally, we calculate the

ensemble-averaged scalar flux by averaging the resulting scalar fluxes over all physical realizations

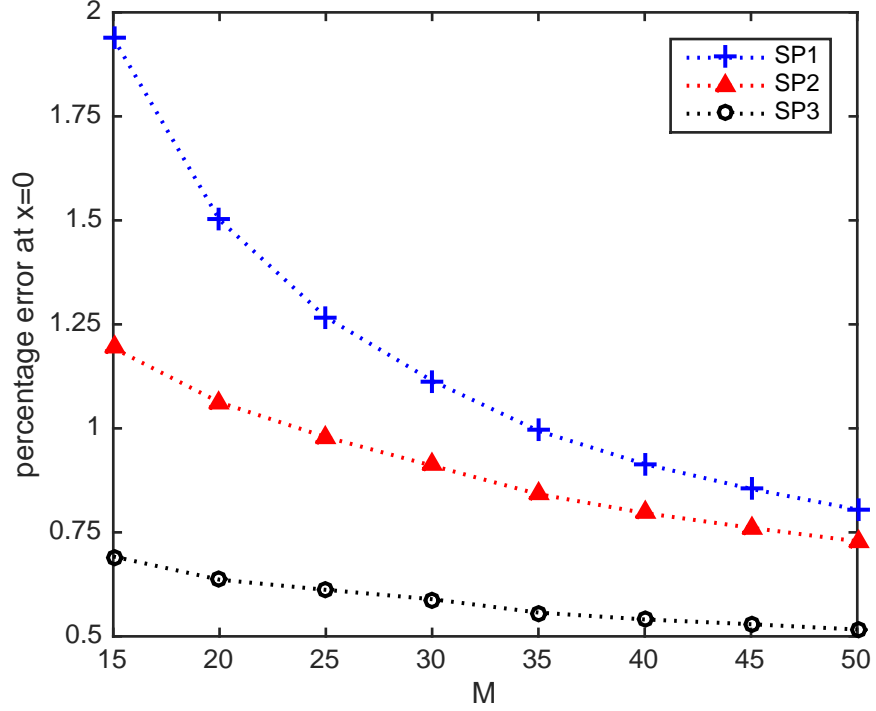


Figure 1: Error of the nonclassical SP_N estimates for the scalar flux with respect to the benchmark solution at $x = 0$.

We calculate the scalar flux at the center of the system ($x = 0$) for different values of M . The expectation is that the estimates obtained with the nonclassical SP_N approximations will increase in accuracy as M increases and $\varepsilon \rightarrow 0$.

Figure 1 shows the percent relative error of the nonclassical SP_N estimates with respect to the benchmark result. As anticipated, the numerical results confirm the asymptotic analysis: (i) the accuracy of the SP_N equations increases as N increases; and (ii) for each choice of N , the error decreases as M increases and the system approaches the diffusive limit.

V Conclusions

This paper derives a set of diffusion approximations to the nonclassical transport equation with isotropic scattering through an asymptotic analysis. These approximations reduce to the simplified P_N equations under the assumption of classical transport, and for that reason are labeled *nonclassical* SP_N equations. Explicit equations are given for nonclassical SP_1 (diffusion), SP_2 , and SP_3 ; higher-order equations can be derived by continuing to follow the same procedure.

Numerical results for a 1-D random periodic system are presented, validating the theoretical predictions. Since the analysis does not yield boundary conditions, these numerical results were obtained by manipulating the nonclassical SP_N equations into a structure that allows us to use classical boundary conditions. This result paves the road to a more complete understanding of the diffusive behavior of the nonclassical transport theory.

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